

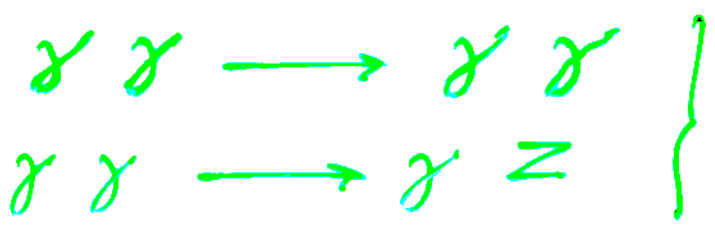
# γγ Collider

Ponyradi  
Renard  
Layrac  
G.

(2)

The case

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Pole  
number of

Old  
VMD

The dominant amplitudes at high energies  
 helicity conserving  
 mainly imaginary

Pomeron  
dominance

## Modern Picture



quarks, leptons, W

$$s_{\gamma\gamma} > (250 \text{ GeV})^2 ; \quad F_{A_1 A_2 A_3 A_4}(s, t, u) = \langle A_3 A_4 | T | A_1 A_2 \rangle$$

- 1) Dominant amplitudes  $\Rightarrow$   
 $F_{++++}(s, t, u), \quad F_{-+-}(s, t, u) = F_{+-+}(s, u, t)$   
 (and their parity partners). Predominantly Imaginary  
 (All other amplitudes)  $\ll \frac{1}{10}$  (Dominant amplitudes)

2)

$W \gg$  quarks, leptons  
 for the dominant amplitudes

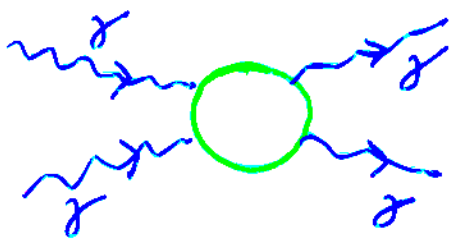
$$\begin{aligned} \gamma\gamma &\rightarrow \gamma\gamma \\ \gamma\gamma &\rightarrow \gamma Z \end{aligned}$$

at  $s_{\gamma\gamma} > (250 \text{ GeV})^2$

\* It is intriguing that

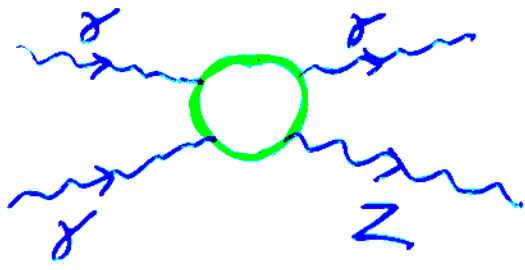
**W-loop** for these processes behaves exactly as expected from **VMD + Pomeron**

\*\* Use them for searching for SUSY and other NP



It depends only on  $Q_\gamma, M_\gamma$

chargino search



It depends only  $Q_{\gamma^+}, M_\gamma, g_V^{\gamma^+}$

\*\*\* Correspondingly for sleptons results depend only on

$$Q_{\tilde{e}}, M_{\tilde{e}}, \epsilon_3^{(\tilde{e})}$$

$$\frac{d\sigma}{dz d\cos\theta^*} = \frac{dL_{\gamma\gamma}}{dz} \left\{ \frac{d\bar{\sigma}_0}{d\cos\theta^*} + \langle \bar{\sigma}_2 \bar{\sigma}_2' \rangle \frac{d\bar{\sigma}_{22}}{d\cos\theta^*} + \langle \bar{\sigma}_3 \bar{\sigma}_3' \rangle \cos(2\phi) \frac{d\bar{\sigma}_3}{d\cos\theta^*} + \langle \bar{\sigma}_3 \bar{\sigma}_3' \rangle \cos(2\phi') \frac{d\bar{\sigma}_3'}{d\cos\theta^*} + \langle \bar{\sigma}_3 \bar{\sigma}_3' \rangle \left[ \cos(2(\phi+\phi')) \frac{d\bar{\sigma}_3}{d\cos\theta^*} + \cos(2(\phi-\phi')) \frac{d\bar{\sigma}_3'}{d\cos\theta^*} \right] + \langle \bar{\sigma}_2 \bar{\sigma}_3' \rangle \sin(2\phi') \frac{d\bar{\sigma}_{23}}{d\cos\theta^*} - \langle \bar{\sigma}_3 \bar{\sigma}_2' \rangle \sin(2\phi) \frac{d\bar{\sigma}_{23}}{d\cos\theta^*} \right\}$$

$q_0 \Rightarrow \sum_{\lambda_3 \lambda_4} \left\{ |F_{++\lambda_3 \lambda_4}|^2 + |F_{+-\lambda_3 \lambda_4}|^2 \right\} \sim |F_{++++}|^2 + |F_{+--+}|^2$   
 $q_{22} \Rightarrow \sum_{\lambda_3 \lambda_4} \left\{ |F_{++\lambda_3 \lambda_4}|^2 - |F_{+-\lambda_3 \lambda_4}|^2 \right\} \sim |F_{++++}|^2 - |F_{+--+}|^2$   
 $q_{02} \Rightarrow -\frac{1}{2} \sum_{\lambda_3 \lambda_4} \text{Re} \left[ F_{++\lambda_3 \lambda_4} F_{-+\lambda_3 \lambda_4}^* \right] \sim$   
 $q_{20} \Rightarrow -\frac{1}{2} \sum_{\lambda_3 \lambda_4} \text{Re} \left[ F_{+-\lambda_3 \lambda_4} F_{++\lambda_3 \lambda_4}^* \right] \sim$   
 $q_{33} \Rightarrow \sum_{\lambda_3 \lambda_4} \text{Re} \left[ F_{+-\lambda_3 \lambda_4} F_{-+\lambda_3 \lambda_4}^* \right] \sim 2 \text{Re} \left\{ F_{+--+} F_{+--+}^* \right\}$   
 $q_{30} \Rightarrow \sum_{\lambda_3 \lambda_4} \text{Re} \left[ F_{++\lambda_3 \lambda_4} F_{-+\lambda_3 \lambda_4}^* \right]$   
 $q_{23} \Rightarrow \sum_{\lambda_3 \lambda_4} \text{Im} \left[ F_{++\lambda_3 \lambda_4} F_{-+\lambda_3 \lambda_4}^* \right]$   
 $q_{23}' \Rightarrow \sum_{\lambda_3 \lambda_4} \text{Im} \left[ F_{++\lambda_3 \lambda_4} F_{-+\lambda_3 \lambda_4}^* \right]$

} similar for  $\gamma Z$

$$\gamma\gamma \rightarrow \gamma\gamma$$

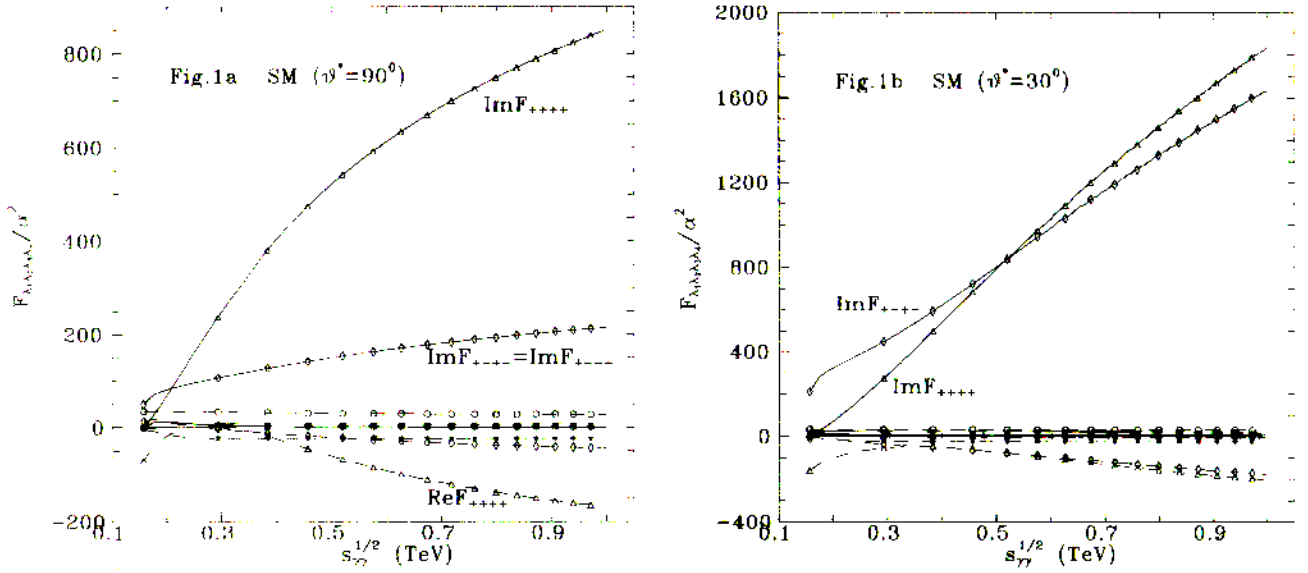
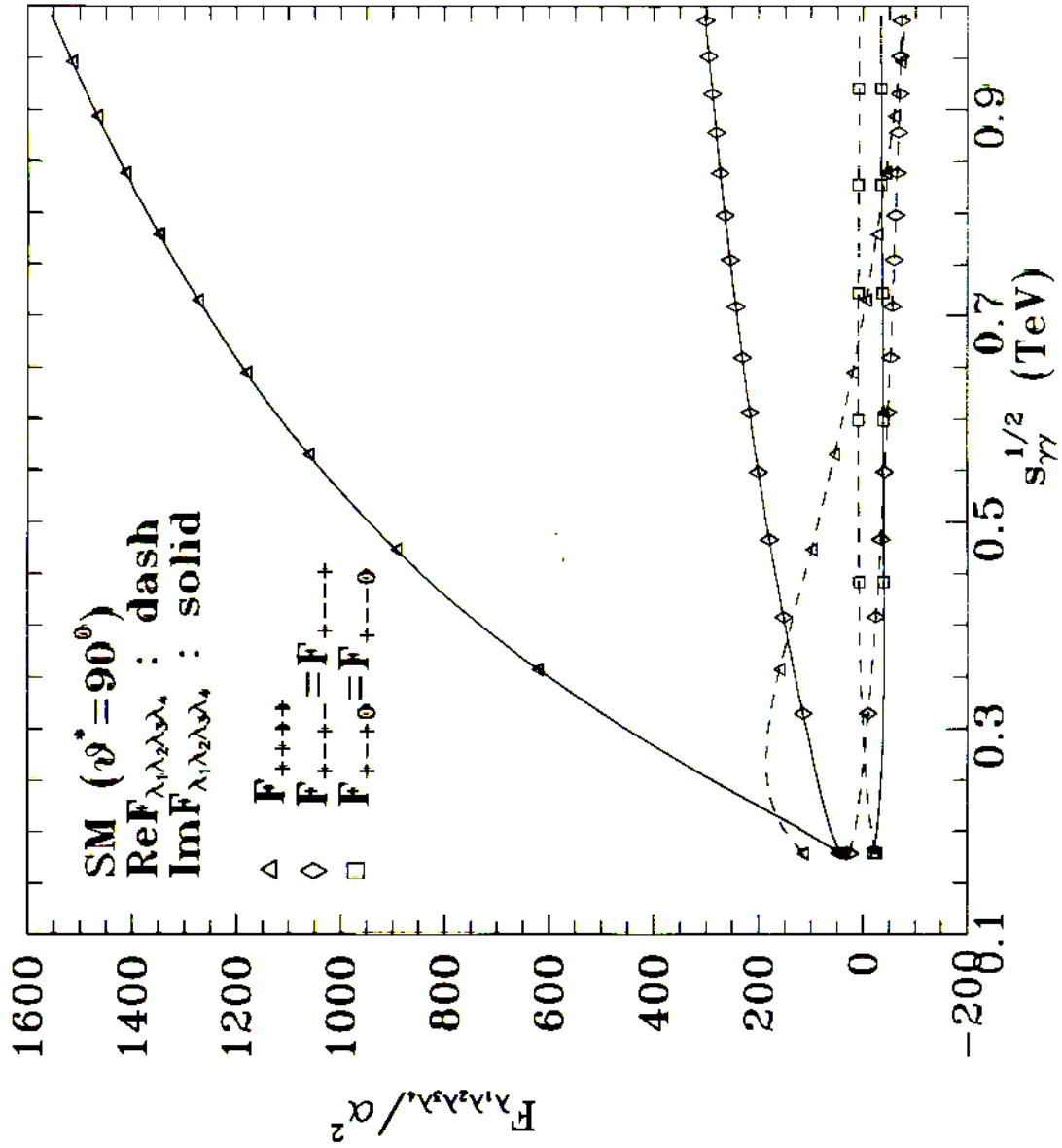


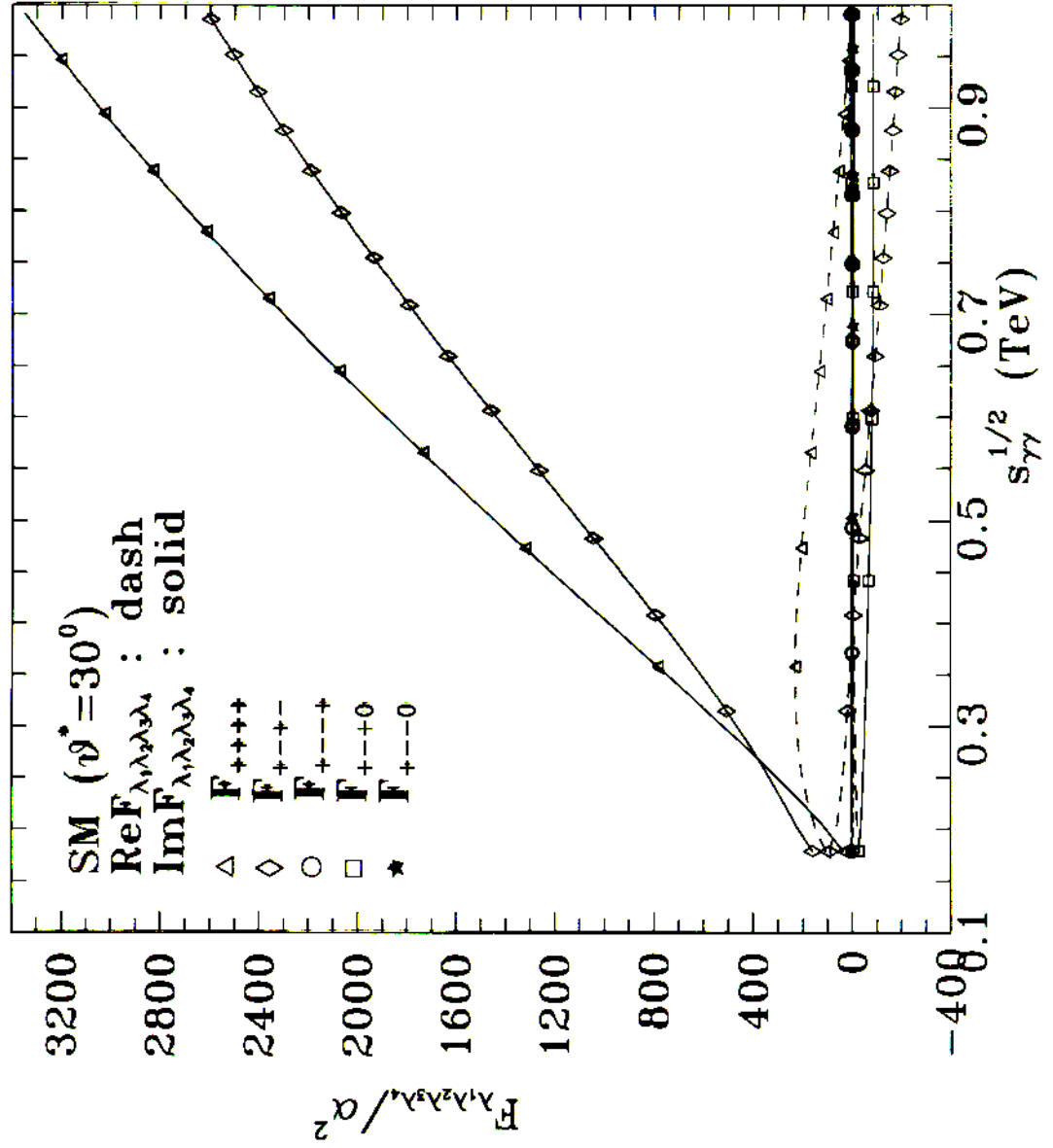
Figure 1: Imaginary (solid line) and real (dashed line) parts of the SM  $\gamma\gamma \rightarrow \gamma\gamma$  helicity amplitudes at  $\vartheta = 90^\circ$  (1a) and  $\vartheta = 30^\circ$  (1b) for  $F_{++++}$  (triangles),  $F_{+--+}$  (circles),  $F_{+---}$  (stars),  $F_{-+-+}$  (rhombs). Finally  $F_{-++-}$  (crosses) is omitted in the (1a) case, since it is identical to  $F_{+--+}$  at  $\vartheta = 90^\circ$ , while in the (1b) case it is negligibly small. We also note that  $F_{+--+}(\hat{s}, \hat{t}, \hat{u}) = F_{-+-+}(\hat{s}, \hat{t}, \hat{u}) = F_{+---}(\hat{s}, \hat{t}, \hat{u}) = F_{-++-}(\hat{s}, \hat{t}, \hat{u})$ .

$\gamma\gamma \rightarrow \gamma Z$



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$\gamma\gamma \rightarrow \gamma Z$



9

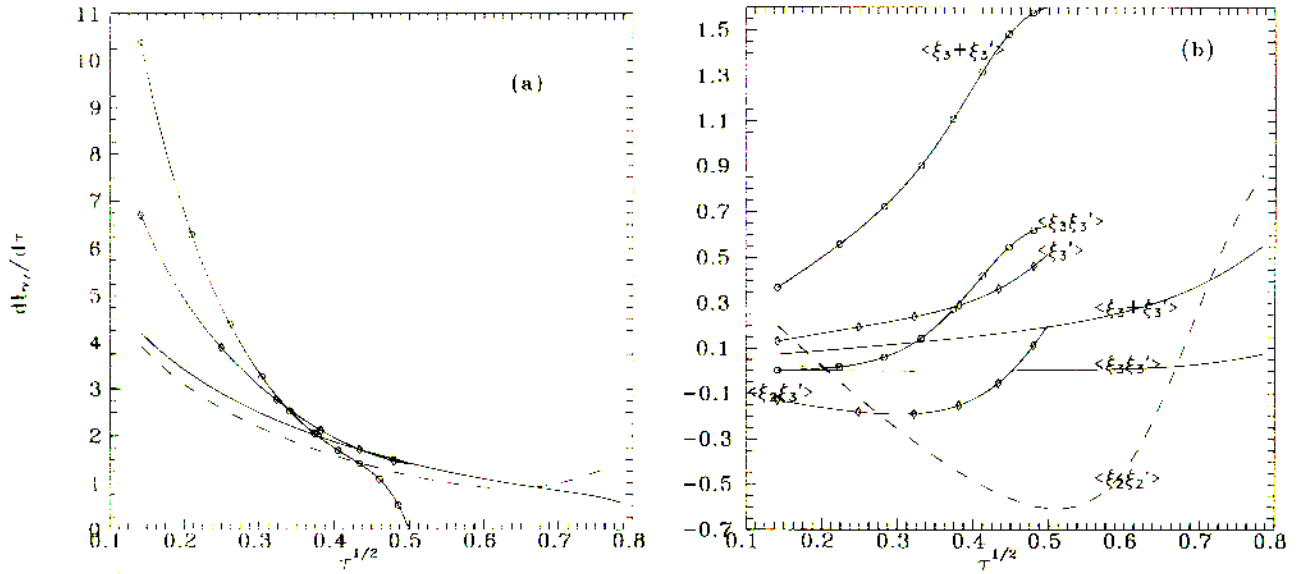


Figure 4: Overall flux factor (a) and elements of the normalized density matrix (b), for the two backscattered photons with  $P_e = P_e' = 0.8$ ,  $P_\gamma = P_\gamma' = -1$ ,  $P_t = P_t' = 0$ ,  $x_0 = x_0' = 4.83$ . (dash):  $P_e = P_e' = P_\gamma = P_\gamma' = 0$ ,  $P_t = P_t' = 1$ , and  $x_0 = x_0' = 4.83$  (solid) or  $x_0 = x_0' = 1$  (circles);  $P_e = 0.8$ ,  $P_\gamma = -1$ ,  $P_t = 0$ ,  $x_0 = 4.83$ ,  $P_e' = 0$ ,  $P_\gamma' = 0$ ,  $P_t' = 1$ ,  $x_0' = 1$ . (rhombs).

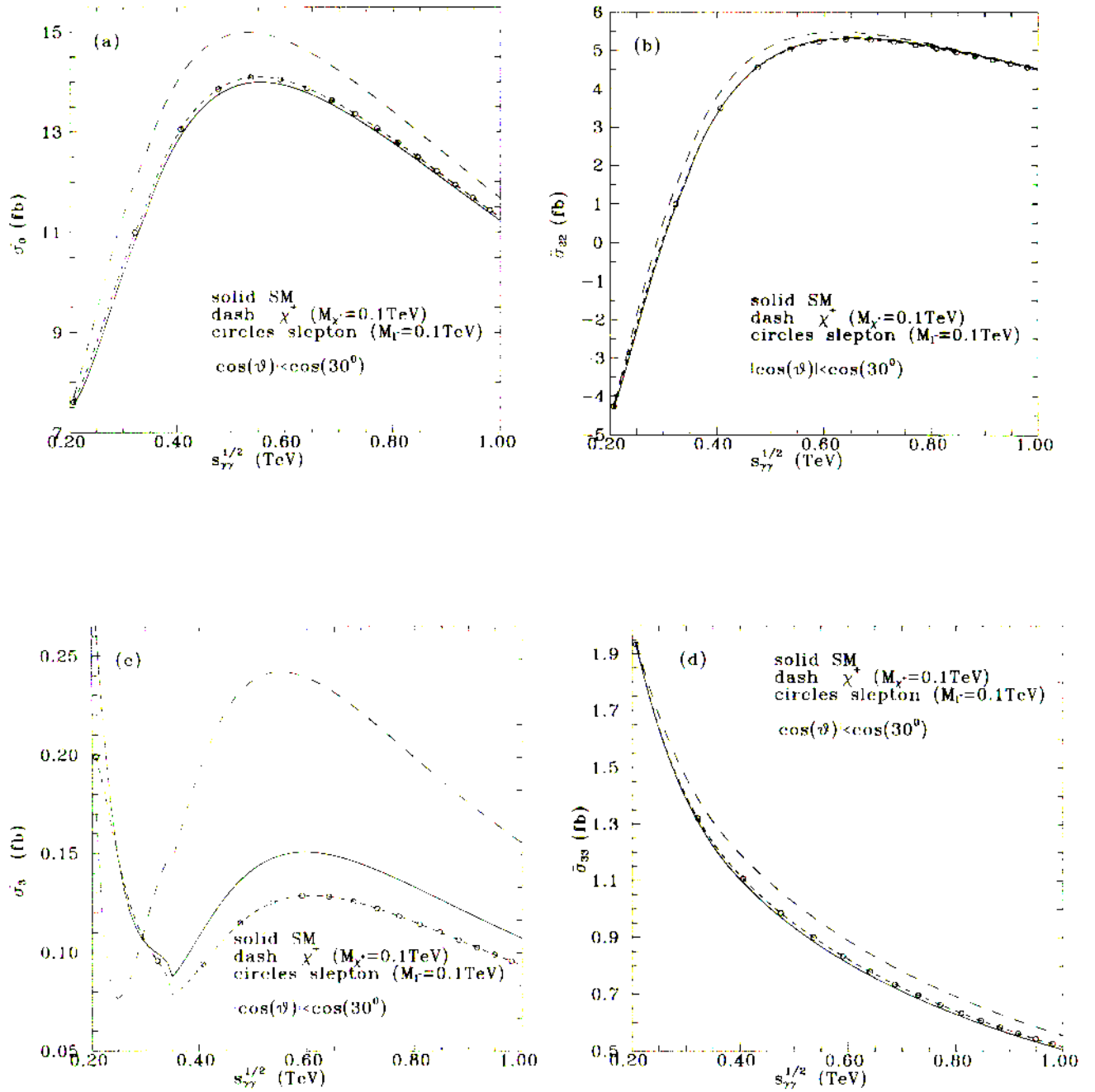


Figure 2:  $\sigma_0$ ,  $\sigma_{22}$ ,  $\sigma_3$  and  $\sigma_{33}$  for SM (solid) and in the presence of a chargino (dash) or a charged slepton (circles) contribution.

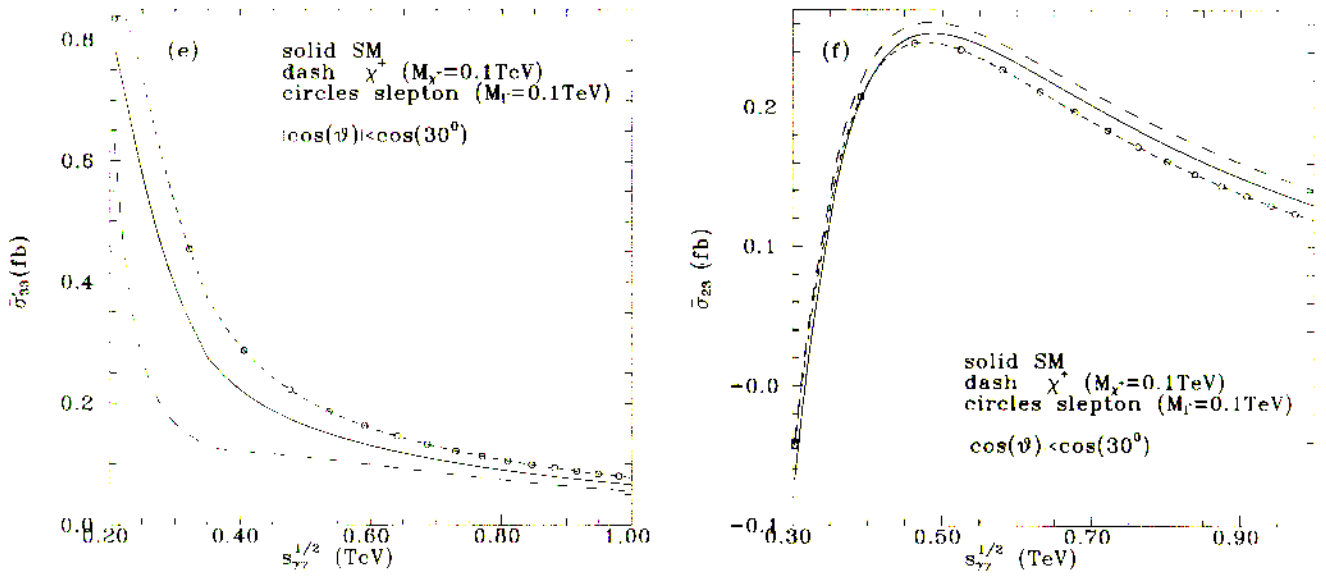
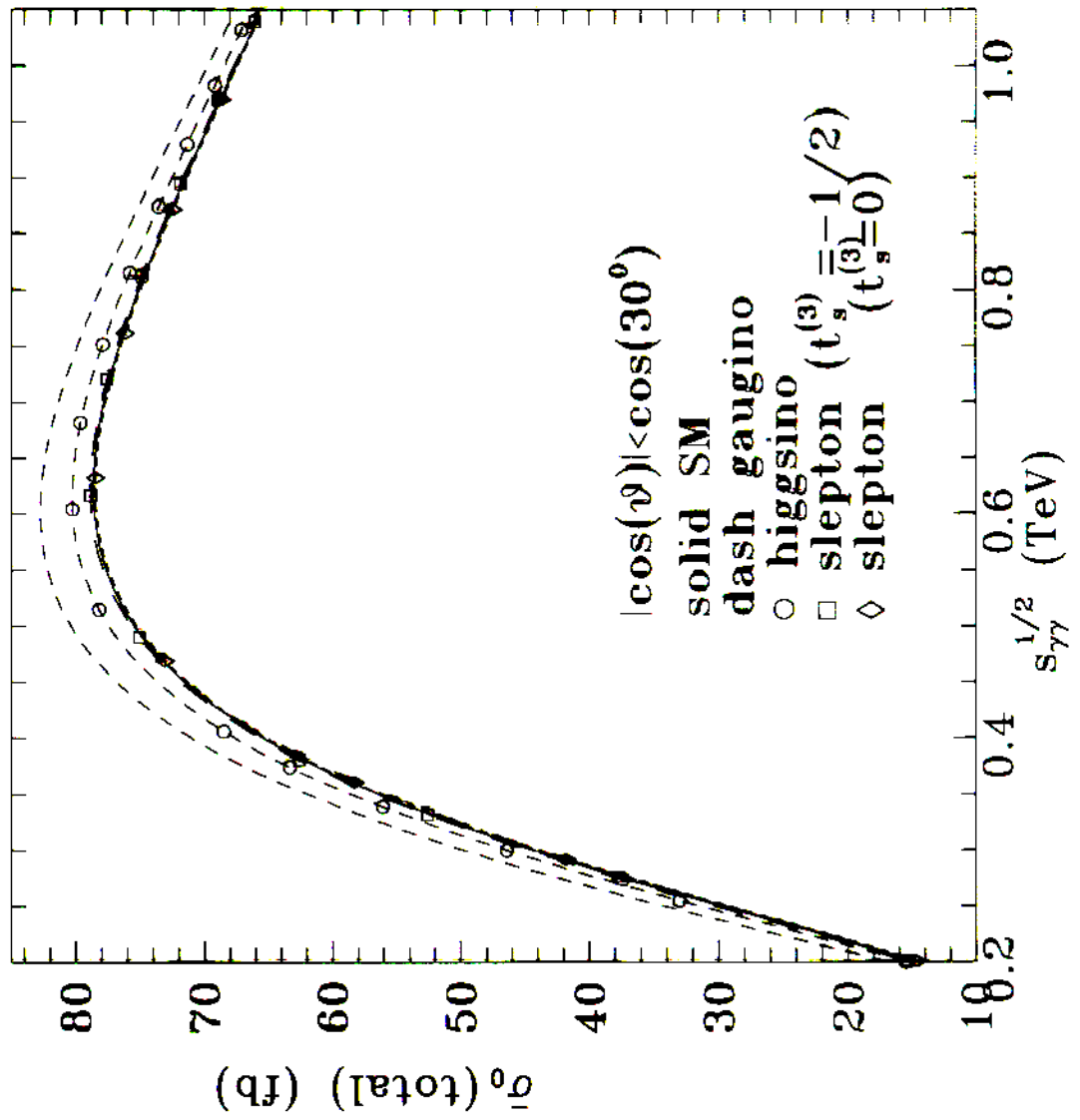


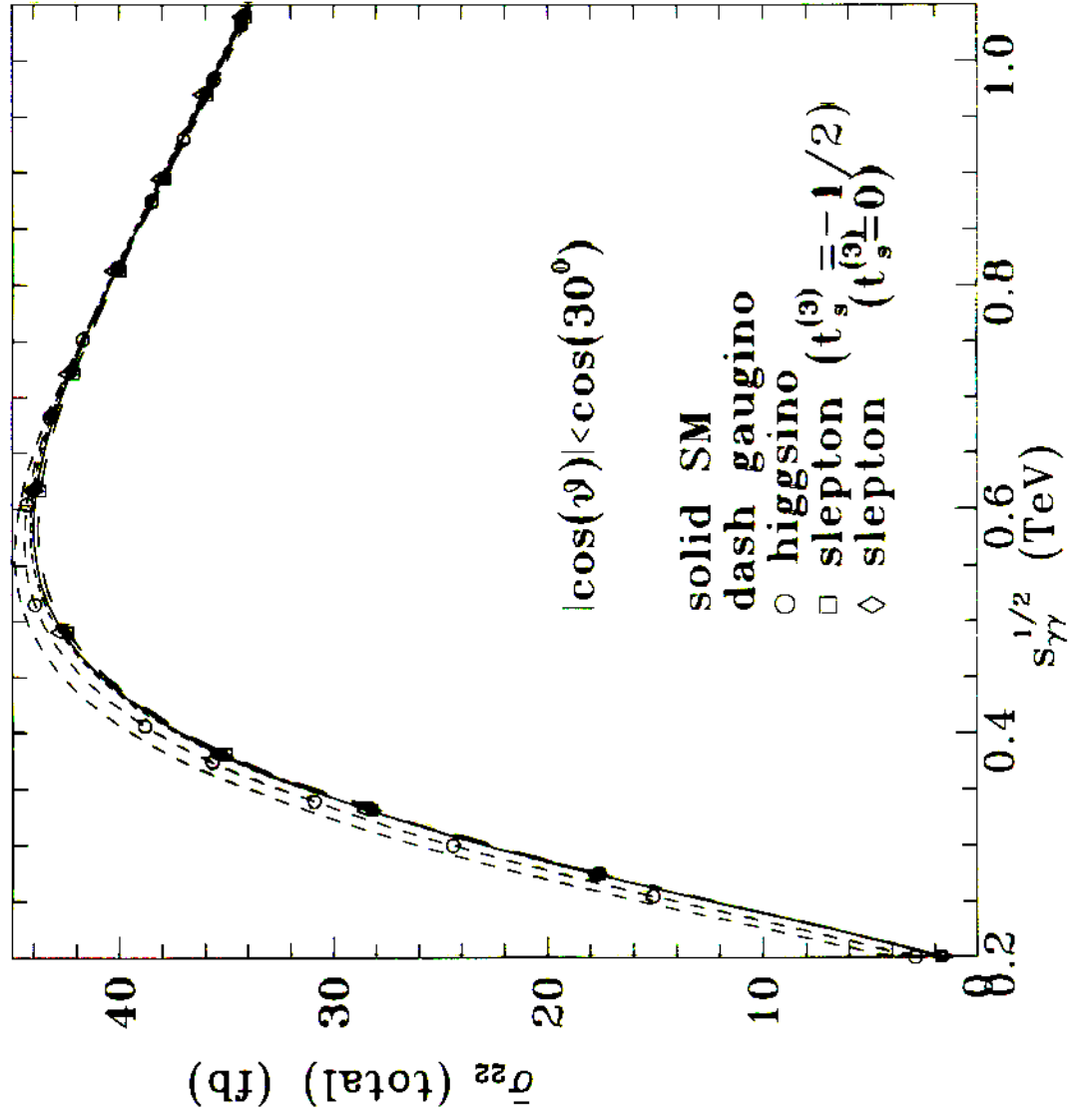
Figure 2:  $\sigma'_{33}$  and  $\sigma_{23}$  for SM (solid) and in the presence of a chargino (dash) or a charged slepton (circles) contribution.

$\gamma\gamma \rightarrow \gamma Z$  (preliminary)



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$\gamma\gamma \rightarrow \gamma Z$  (preliminary)



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# conclusion

(12)

1) Chargino effects in  $\bar{\sigma}_0 (\gamma\gamma \rightarrow \gamma\gamma)$

*If only  $\gamma\gamma \rightarrow \gamma\gamma$  is used*

3-4 SD  $\longleftrightarrow M_{\chi^\pm} = 200 \text{ GeV}$

1 SD  $\longleftrightarrow M_{\chi} \approx 250 \text{ GeV}$

2) Effect stronger for  $\gamma\gamma \rightarrow \gamma Z$

3) Combine them finally

If more SUSY particles exist their effect to  $\bar{\sigma}_0$  always cumulative (positive)

for both

$\gamma\gamma \rightarrow \gamma\gamma$

$\gamma\gamma \rightarrow \gamma Z$